PHOTOELECTRIC PHOTOMETRY OF ALGOL AND HAMAL STARS

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Final Year Project Report Submitted in Partial Fulfilment of the Requirements for the Degree of Bachelor of Science (Hons) Physics in the Faculty of Applied Sciences Universiti Teknologi MARA

NOVEMBER 2010

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ACKNOWLEDGEMENTS

Upon completion of this project, I would like to express my gratitude to many parties.

First of all, I would like to express my gratitude to Allah S.W.T, because of His

blessings; this thesis was completed on time. I would like to express my deepest

appreciation to my Supervisor, Miss Zety Sharizat Binti Hamidi because of his valuable

comments, suggestions, advice, precious ideas and guidance during the process of

completing this thesis. I would like to thank to lecturers and students from University

Malaya, who's always keep helping me during laboratory work and give me a lot of

information in order to understand the result of my project. In addition, I would like to

express my special thanks to my friends Husna, Fatimatun, Adibah and Azlina, for their

advice and support throughout the completion of this thesis. Special dedications to my

beloved parents, Mohd Noris Bin Mansor and Fauziyah Binti Rozali, for their

understanding and support. Last but not least, thank to Faculty of Applied Science,

Universiti Teknologi Mara for give me opportunity to do final year project.

Nazirah Binti Mohd Noris

(2008288048)

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ABSTRACT

PHOTOELECTRIC PHOTOMETRY OF ALGOL AND HAMAL STARS

Photoelectric Photometry is technique of astronomy concerned with measuring flux, the most fundamental and oldest research technique and gives information about the total energy emitted by the object, its size, its temperature and other physical properties. Algol star is variable star that changes the brightness of the star. From the surface temperature range 10,000K until 28,000K is B spectral class, appear more bluish colour and have light curve. The comparison star is Hamal star is non- variable star that not changes brightness of the star. Hamal has the surface temperature range from 3500K to 4900K so it can classified in K spectral and the colour index (B-V) is positive value, so the star appear reddish colour.